

An identity for R_n embedded into E_{n+1}

GUILLERMO GONZÁLEZ, JOSÉ L LÓPEZ-BONILLA and
MARCO A ROSALES*

Area de Física, División de CBI, UAM-Azcapotzalco, Apdo. postal 16-306, 02200 México D.F., Mexico

*Depto. de Física y Matemáticas, Escuela de Ciencias, Universidad de las Américas Puebla, Apdo. Postal 100, Sta. Catarina Mártir, 72820 Puebla, Mexico

MS received 15 November 1993

Abstract. An identity is obtained for a Riemannian n -space (R_n) locally and isometrically embedded into a pseudo-Euclidean $(n + 1)$ -space (E_{n+1}), relating the corresponding second fundamental form with the intrinsic geometry of R_n . For $n = 4$ such an identity reduces to a previous result by Goenner.

Keywords. Riemannian spaces embedding; local and isometric embedding.

PACS Nos 04-20; 04-90

1. Introduction

We will say that R_n is locally and isometrically embedded into E_{n+1} if and only if there exists the second fundamental form tensor $b_{ac} = b_{ca}$ that satisfies the equations [1]

$$R_{ijkc} = \varepsilon(b_{ik}b_{jc} - b_{ic}b_{jk}) \quad \text{Gauss} \quad (1a)$$

and

$$b_{ij;k} = b_{ik;j} \quad \text{Codazzi} \quad (1b)$$

where R_{ajqc} is the curvature tensor for R_n , $\varepsilon = \pm 1$ is the indicator for the normal to this n -space and; c denotes the covariant derivative.

In a usual embedding problem what is known is the R_n intrinsic geometry built from the metric tensor $g_{ab} = g_{ba}$, and the aim is to find out if the tensor b_{ac} satisfying (1a) and (1b) exists. For example, if $n = 4$ and the space-time has a vanishing Ricci tensor, then we know [1–3] that this b_{ac} does not exist and, hence, it may be concluded that no empty R_4 accepts embedding into E_5 . Nevertheless, there are other metrics for which such tensor does exist [1, 4–7], and it is in these cases where a relation or identity to explicitly build the b_{ac} is needed.

In §2 we will show that, using the Gauss equation (1a) only, it is possible to obtain an identity connecting b_{ac} with the intrinsic geometry of R_n , and we will indicate the conditions under which such an identity allows the explicit derivation of the second fundamental form. When $n = 4$ this identity reduces to a relation previously published by Goenner [8] as will be proved in §3.

2. b_{ac} and the intrinsic geometry of R_n

The Ricci tensor:

$$R_{jk} \equiv R_{jki}^i = \varepsilon(b^i_k b_{ji} - bb_{jk}), \quad b \equiv b^r_r \tag{2a}$$

and the scalar curvature:

$$R \equiv R^j_j = \varepsilon(b^{ac} b_{ac} - b^2) \tag{2b}$$

are obtained with the help of the Gauss equation. Then, (1a), (2a) and (2b) lead to the identity:

$$pb_{ac} = \frac{R}{2} R_{ac} + \frac{1}{2} R_a^r R_{rc} - R_{rc} - R_{arqc} R^{rq} - \frac{1}{4} R_{arqt} R_c^{rqt} \tag{3a}$$

that has not been found in the literature thus far. In this equation:

$$p \equiv b^{ra} \left(b^i_r b_{ia} - \frac{3}{2} bb_{ra} + \frac{b^2}{2} g_{ra} \right). \tag{3b}$$

But the Einstein tensor is given by:

$$G_{ra} = R_{ra} - \frac{R}{2} g_{ra} \tag{3c}$$

and, hence, (2a), (2b), (3b) and (3c) imply that

$$p = \varepsilon b^{ra} G_{ra} \tag{4a}$$

so that, after multiplying (3a) by εG^{ac} we obtain:

$$p^2 = \varepsilon \left(\frac{1}{2} R_a^r R_{rc} - R_{arqc} G^{ra} - \frac{1}{4} R_{arqt} R_c^{rqt} \right) G^{ac} \geq 0 \tag{4b}$$

which determines the value of ε when $p \neq 0$. From (4b) we learn that p is a function of the intrinsic geometry of R_n only.

Therefore, we can give to the embedding process the following sequence:

- 1) We have g_{ac} and do not know whether this R_n accepts a local and isometric embedding into E_{n+1} .
- 2) We compute p through (4b). Then we will have one of the two following cases:
 - a) If $p \neq 0$ then we determine b_{ac} with the help of (3a). Hence, the R_n does accept embedding into E_{n+1} if and only if *this* b_{ac} fulfils the Codazzi condition, eq. (1b).
 - b) When $p = 0$ eq. (3a) does not provide any explicit information about b_{ac} . Then, in order to decide if R_n is of class one, and to construct the corresponding second fundamental form, it is necessary to solve (1a) and (1b) directly.

3. The case $n = 4$

For ordinary space-time we can simplify (3a) and (4b) by making the dimension 4 more explicit. Indeed, recalling that the conformal tensor C_{arqt} and the double dual

An identity for R_n embedded into E_{n+1}

tensor $*R^*_{arqt}$ are related through the expression [9].

$$R_{arqt} = *R^*_{arqt} + 2C_{arqt} - \frac{R}{6}(g_{aq}g_{rt} - g_{at}g_{rq}) \quad (5a)$$

this can be introduced into the last term of (3a) to yield

$$pb_{ac} = \frac{1}{4} \left(\frac{K_2}{4} + R^2 - 3R_{qt}R^{qt} \right) g_{ac} - \frac{3}{2} C_{aqt} R^{qt} + \frac{3}{2} \left(R_{aq} - \frac{R}{6} g_{aq} \right) R^q_c \quad (5b)$$

where Lanczos identity [10]

$$*R^{*a}_{cqt} R^{rcqt} = \frac{K}{4} g^{ar}, \quad K_2 \equiv *R^{*ijk} R_{ijk} \quad (5c)$$

has been used.

Equation (5b) is valid only for $n = 4$ and it is equivalent to (3.1) of Goenner (1976). Furthermore, (4b) reduces to

$$p^2 = -\frac{3\varepsilon}{2} \left(\frac{R}{24} K_2 + R_{iqj} G^{ij} G^{qt} \right) \geq 0. \quad (5d)$$

As an application of (5b) and (5d) consider Gödel [11] cosmological model (signature + 2):

$$ds^2 = -(dx^1)^2 - 2e^{x^4} dx^1 dx^2 - \frac{1}{2} e^{2x^4} (dx^2)^2 + (dx^3)^2 + (dx^4)^2 \quad (6a)$$

which, together with (5d) imply that $\varepsilon = 1$ and $p = 3\sqrt{2}/4$. Therefore, from (5b) we get $b_{ac} = 0$, except

$$b_{11} = b_{44} = -\frac{\sqrt{2}}{2}, \quad b_{12} = -\frac{\sqrt{2}}{2}, \quad b_{22} = -\frac{3\sqrt{2}}{4} e^{2x^4} \quad (6b)$$

but this second fundamental form tensor does not satisfy (1b) since, for example, $b_{12:4} \neq b_{14:2}$. Then, due to remark 2 a) at the end of last section, we conclude that Gödel solution does not accept local and isometric embedding into E_5 as was already shown by Szekeres [2] using another technique. It is still ignored if (6a) admits embedding into E_6 (see [6]).

Notice that, when $p = 0$, it is not possible to use (5b) to construct b_{ij} and it is necessary to analyze the Gauss and Codazzi equations directly. An example of such cases can be found in Collinson [5], where Einstein-Maxwell metrics embedded into E_5 are studied.

Acknowledgements

The authors express their gratitude to Dr E Piña-Garza for suggestions that led to identity (3a).

References

- [1] D Kramer, H Stephani, M Mac Callum and E Herlt *Exact solutions of Einstein's field equations* (Cambridge: University Press, 1980)
- [2] Szekeres, *Nuovo Cimento* **A43**, 1062 (1966)
- [3] R Fuentes, J L López-Bonilla, Ovando G and T Matos, *Gen. Relativ. Gravit.* **21**, 777 (1989)
- [4] J Rosen, *Rev. Mod. Phys.* **37**, 204 (1965)
- [5] C D Collinson, *Commun. Math. Phys.* **8**, 1 (1968a)
- [6] C D Collinson, *J. Math. Phys.* **9** 403 (1968b)
- [7] A Barnes, *Gen. Relativ. Gravit.* **5**, 147 (1974)
- [8] H F Goenner, *Tensor N. S.* **30**, 15 (1976)
- [9] C Lanczos, *Rev. Mod. Phys.* **34**, 379 (1962)
- [10] C Lanczos, *Ann. Math.* **39**, 842 (1938)
- [11] K Gödel, *Rev. Mod. Phys.* **21**, 447 (1949)