

# 2-D Crustal thermal structure along Thuadara–Sindad DSS profile across Narmada–Son lineament, central India

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Central India is traversed by a WSW–ENE trending Narmada–Son lineament (NSL) which is characterized by the presence of numerous hot springs, feeder dykes for Deccan Traps and seismicity all along its length. It is divided in two parts by the Barwani–Sukta Fault (BSF). To the west of this fault a graben exists, whereas to the east the basement is uplifted between Narmada North Fault (NNF) and Narmada South Fault (NSF). The present work deals with the 2-D thermal modeling to delineate the crustal thermal structure of the western part of NSL region along the Thuadara–Sindad Deep Seismic Sounding (DSS) profile which runs almost in the N–S direction across the NSL. Numerical results of the model reveal that the conductive surface heat flow value in the region under consideration varies between 45 and 47 mW/m<sup>2</sup>. Out of which 23 mW/m<sup>2</sup> is the contribution from the mantle heat flow and the remaining from within the crust. The Curie depth is found to vary between 46 and 47 km and is in close agreement with the earlier reported Curie depth estimated from the analysis of MAGSAT data. The Moho temperature varies between 470 and 500°C. This study suggests that this western part of central Indian region is characterized by low mantle heat flow which in turn makes the lower crust brittle and amenable to the occurrence of deep focused earthquakes such as Satpura (1938) earthquake.

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## 1. Introduction

The NSL is a well-defined geological feature in central India extending about 1600 km from Bharuch on the western coast towards the NE underneath the Monghyr–Saharsa ridge between 72 to 88°E and 21°30' to 24°N. Mishra (1977) has suggested the possible extension of NSL westward into the Arabian Sea and eastward up to the Shillong plateau. It is believed to have originated during the middle to late Archaean period and has influenced the deposition of Neoproterozoic Vindhyan sediments to its north and Gondwana sediments to its south (West 1962; Radhakrishna 1989). The unusual features of NSL compared to other parts of the Indian shield are the occurrence of deeper earthquakes, namely the 1938

Satpura earthquake (21°32'N, 75°50'E) and the 1997 Jabalpur earthquake (23°5'N, 80°2'E) with focal depths (> 35 km) in the lower crust while focal depths of most of Indian shield earthquakes are confined to 10 km (Mukherjee 1942; Rao *et al* 2002; Rajendran and Rajendran 1998; Gahalaut *et al* 2004). Several geological and geophysical studies have been carried out to understand the evolution of NSL and its influence on tectonic framework of central India. These studies include four Deep Seismic Sounding (DSS) profiles namely (I) Hrapur–Mandla, (II) Khajuriakalan–Pulgaon, (III) Ujjain–Mahan, and (IV) Thuadara–Sindad (Kumar *et al* 2000; Tewari *et al* 2001; Sridhar and Tewari 2001; Kumar 2002; Mall *et al* 2002). Locations of these profiles are shown in figure 1. From the analysis of geological and geophysical data, it has been

**Keywords.** 2-D thermal modeling; surface heat flow; crustal thermal structure; Curie depth, Narmada–Son lineament; Moho temperature.

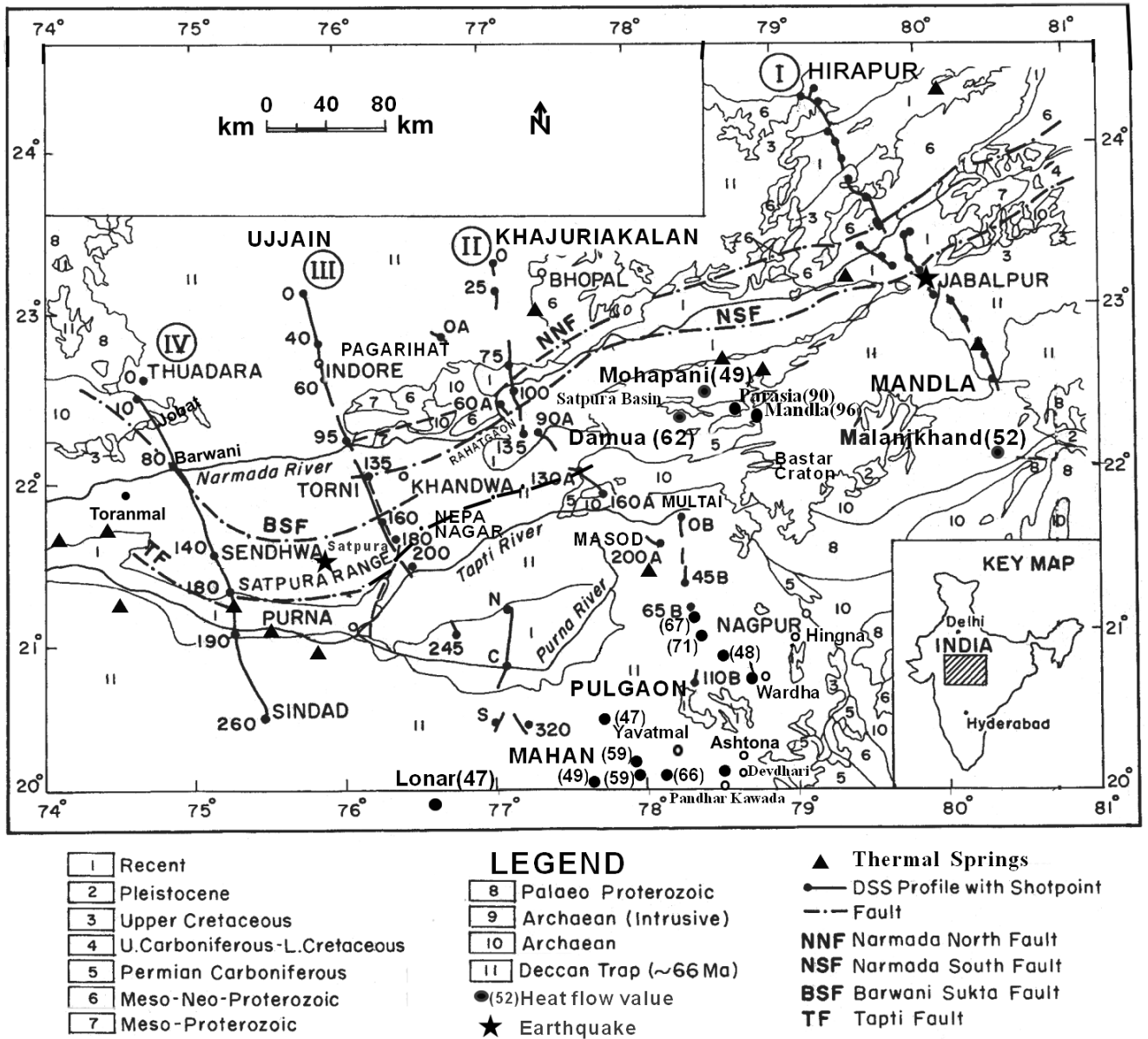


Figure 1. Geological and tectonic map of Narmada–Son lineaments with locations of DSS profiles and heat flow sites (modified after Sridhar and Tewari 2001).

suggested that the NSL is divided by a curvilinear Barwani–Sukta fault which is marked by shearing and brecciation of basalt, alignment of hot springs, and presence of earthquake epicenters (Sridhar 2001; Tewari *et al* 2002). The southwestern part represents a graben and the northeastern part a basement uplift bounded by deep-seated Narmada North Fault (NNF) and Narmada South Fault (NSF) (figure 1). The first three DSS profiles traverse the region of basement uplift while the last one profile traverses the region of graben.

This paper deals with 2-D thermal modeling along the 260 km long Thuadara–Sindad DSS profile. The 2-D modeling approach needs 2-D crustal structure up to the Moho and  $P$ -wave velocity distribution for construction of the model and for

estimation of heat production in the middle and lower crust respectively, by using an empirical relationship between heat production ( $A$ ) and  $P$ -wave velocity ( $V_p$ ).

## 2. Geology and tectonic framework

The region of Thuadara–Sindad DSS profile is mostly covered by Deccan Traps. The Deccan Volcanic rocks are predominantly Fe-rich tholeiites formed at the Cretaceous Tertiary boundary (~65 Ma ago) (Courtilot 1994; Mahoney *et al* 2000). Other geological formations exposed in the northern part include marine Cretaceous deposits belonging to the Bagh group, Lameta beds

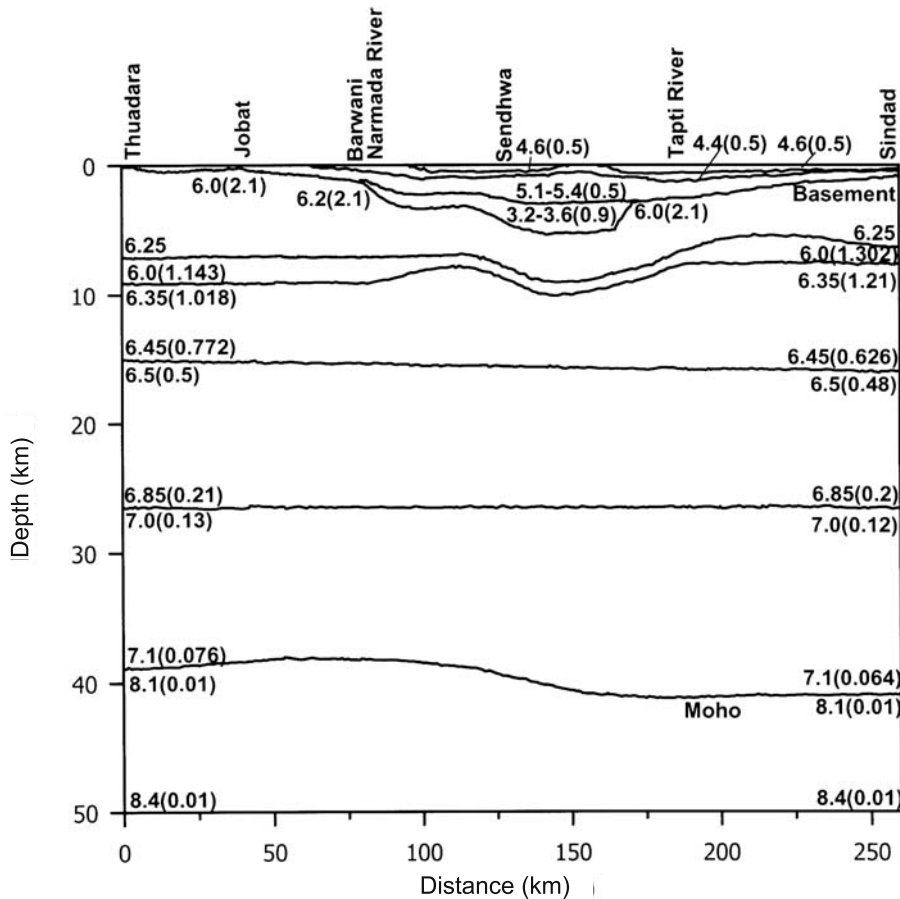


Figure 2. Crustal layers with  $V_p$ -velocity distribution (km/s) and distribution of heat generation ( $\mu\text{W}/\text{m}^3$ ) values (in brackets) along the profile.

(Mesozoics), the Aravalli and granitic gneisses. The Bagh fauna has close affinities with that of the Cretaceous fauna of Gujarat in western India. So it can be presumed that the lower reaches of the Narmada River was a low depression between Satpuras and the Vindhya during Bagh times and an arm of the sea encroached the Narmada Son lineament in the form of a marine transgression (Sridhar 2001). The southern part of Thuadara–Sindad profile is traversed by an E–W trending Tapti fault which marks the southern boundary of the Satpura ranges and northern boundary of Tapti Alluvium. The Satpura range between the Narmada and Tapti Rivers is considered as a horst based on broad gravity high (Qureshy 1964). The Satpura horst is bounded by Son–Narmada and Tapti faults in the north and south, respectively. These faults have given rise to two rift valleys on both sides of the horst. The Narmada and Tapti Rivers flow in these rift valleys. For completeness of the work a brief description of the crustal structure and  $P$ -wave velocity distribution along the Thuadara–Sindad profile, given by Sridhar (2001); Sridhar and Tewari (2001); and Tewari and Kumar (2003) is presented in the next section.

### 3. Crustal structure and $P$ -wave velocity distribution

The crustal structure and  $P$ -wave velocity distribution along the Thuadara–Sindad profile is shown in figure 2. The interpretation of seismic data along this profile reveals a sedimentary graben between the Narmada and Tapti Rivers (Sridhar 2001; Sridhar and Tewari 2001). Within the graben the depth to the basement ( $6.0$ – $6.2 \text{ km s}^{-1}$ ) is about 3000 m between the Narmada River and Sendhwa and 5000–5500 m between the Sendhwa and Tapti River. This graben contains 1000–2800 m thick low velocity (velocity  $3.2$ – $3.6 \text{ km s}^{-1}$ ) sediments under a thick cover (max  $\sim 2500$  m) of the Deccan Traps ( $4.4$ – $5.4 \text{ km s}^{-1}$ ). The graben is bounded by the Barwani–Sukta fault in the north and Tapti fault in the south (figure 1). Below the basement layer lies a low velocity ( $6.0 \text{ km s}^{-1}$ ) zone of  $\sim 2$  km thickness. Below this layer two more layers with velocity in the range of  $6.35$ – $6.45 \text{ km s}^{-1}$  and  $6.5$ – $6.85 \text{ km s}^{-1}$  have been identified. A high velocity ( $7.0$ – $7.1 \text{ km s}^{-1}$ ) mafic underplating layer exists at a depth of about 27 km. Its lower boundary forms the Moho depth which varies between 38 and 42 km.

#### 4. Mathematical modeling

Conduction is the most dominant process of heat transfer in solid materials. Hence, this mode of heat transfer is generally considered in crustal thermal modeling. A finite element based numerical program named Numerically Integrated System Analysis (NISA), is used for computation of the 2-D crustal thermal structure by solving the 2-D steady state heat conduction equation (Rai and Thiagarajan 2006). For modeling purpose the entire area of the model is divided into a network of four nodes quadrilateral elements with 10 km length in X-direction. Because of the more inhomogeneous character of the upper crustal layer, the region between ground surface and 10 km depths are divided into relatively smaller elements. The width of the elements in the Z-direction is 500 m for the first 10 km while it is 2 km below it. The depth of the lower boundary of the model is taken at 50 km which is below the Moho. One should know distribution of heat generation ( $A$ ), and thermal conductivity ( $K$ ) within the region to be modeled for computation of temperature distribution.

With the precise knowledge of  $P$ -wave velocity distribution available from the DSS studies the distribution of heat generation  $A$ , in the middle and lower crustal layer is estimated from  $P$ -wave velocity value by using the following empirical relationship between  $P$ -wave seismic velocity ( $V_p$ ) and radiogenic heat generation ( $A$ ) for Precambrian rocks (Rybach and Buntbarth 1984; Cermak 1989):

$$\ln A = 12.6 - 2.17 V_p, \quad (1)$$

where  $A$  is in  $\mu\text{W}/\text{m}^3$  and  $V_p$  (20, 100) is the velocity (km/s) at room temperature  $20^\circ\text{C}$  and pressure 100 MPa. The estimated value of  $A$  is corrected for *in-situ* pressure and temperature conditions by using a correction factor as suggested by Cermak (1989). However, this conversion relationship between  $V_p$  and  $A(X, Z)$  is not applicable in the upper crustal layer because the upper crust is highly heterogeneous due to the presence of micro cracks which facilitate the redistribution of radioactive elements by groundwater movement. This may have considerably altered the original distribution of the radioactive content of rocks (Cermak 1989). The heat production values of the top layers of Deccan volcanic and sedimentary rocks up to the basement are taken as constant. For the NSL region, heat production for sedimentary layer ( $3.2\text{--}3.6 \text{ km s}^{-1}$ ) is taken equal to  $0.9 \mu\text{W}/\text{m}^3$  which is the mean of heat production of sandstone, clay and limestone (Rybach and Cermak 1982). Concentrations of U, Th, and K in the Basaltic volcanic rock of the region have been

reported by Chandrasekaram *et al* 1999; Mahoney *et al* 2000 and Sheth *et al* 2004. In this paper, the heat generation value equal to  $0.5 \mu\text{W}/\text{m}^3$  for Deccan volcanic layer ( $4.4\text{--}5.4 \text{ km s}^{-1}$ ) is taken as the mean of heat generation values estimated at 210 m and 1100 m depths from U, Th, K concentration values given by Mahoney *et al* (2000) for tholeiitic flood basalts from Toranmal ( $21^\circ 53' \text{N}$ ,  $74^\circ 28' \text{E}$ ) area located near the profile on its western side. The heat generation value is estimated by using the following conversion factor (Birch 1954):

$$\begin{aligned} \text{HP}(\mu\text{W}/\text{m}^3) \\ = \rho(0.035 C_K + 0.097 C_U + 0.026 C_{\text{Th}}), \quad (2) \end{aligned}$$

where  $\rho$  is the density of the rocks in  $\text{g cm}^{-3}$ ;  $C_K$ ,  $C_U$ , and  $C_{\text{Th}}$  refer to K, U, and Th concentrations by weight (K in %, U and Th in ppm), the numerical constant 0.035 refers to heat production ( $10^{-12} \text{ W}$ ) per gram of rock per 1% of K, and the constants 0.097 and 0.026 refer to heat production ( $10^{-12} \text{ W}$ ) per gram of rock, per 1 ppm of U and 1 ppm of Th, respectively. The density of rock,  $\rho$ , is taken equal to  $2.8 \text{ g m}^{-3}$  for Deccan volcanic rocks (Ray *et al* 2007).

The heat generation value between upper and lower boundaries of basement rocks is estimated by using an exponential model of radiogenic heat source which is defined by (Lachenbruch 1970).

$$A(Z) = A_0 \exp(-Z/D), \quad (3)$$

in which  $A_0$  is the heat production of the basement rock, and  $D$  is the logarithmic decrement which represents thickness of the radioactive enriched crustal layer. In this case  $a < Z < b$  km where  $a$  and  $b$  are the depth of upper and lower boundaries of the basement layer.  $A_0$  and  $D$  values for the NSL are not available. However, Gupta *et al* (1993) have reported these values for the granodiorite rock from Malanjhand ( $22.026^\circ \text{N}$ ,  $80.715^\circ \text{E}$ ) located south of NSL in Bastar craton (figure 1). These values are:  $A_0 = 2.1 \mu\text{W}/\text{m}^3$ ,  $D = 11.5 \text{ km}$  and  $K_0 = 3.1 \text{ W m}^{-1} \text{ K}^{-1}$ . These values for  $A_0$ ,  $K_0$  and  $D$  have been used in this study to also estimate the heat generation distribution from the top of the basement rock to the top of the LVZ. Gupta *et al* (1993) have also given mantle heat flow value equal to  $23 \text{ mW}/\text{m}^2$  for the Bastar craton and the same value has been considered in the present study at the base of the model as a boundary condition. Since heat production in the upper mantle is very small, a small value of  $A$  equal to  $0.01 \mu\text{W}/\text{m}^3$  is assigned to the region beneath the Moho (Safanda *et al* 1992). The calculated values of  $A$  are assigned

Table 1. *Temperature dependent thermal conductivity values.*

	$K_0(\text{Wm}^{-1} \text{K}^{-1})$	$T(^{\circ}\text{C})$	$C(\text{K}^{-1})$
Upper crust	3.1	< 300	0.001
Lower crust	2.0	300–500	0.0
Lithospheric mantle	2.5	> 500	–0.00025

within the model area according to the distribution of corresponding  $V_p$  velocities. Distributions of estimated heat generation values are shown within the brackets in figure 2.

In this study thermal conductivity has been considered as a function of temperature defined by Cermak and Bodri (1986); Shengbiao and Jiyang (2000):

$$K = \frac{K_0}{(1 + CT)}, \quad (4)$$

in which  $K_0$  is the thermal conductivity at surface condition and  $C$  is an experimentally determined constant which controls the behaviour of  $K$  with temperature  $T$ . The values of  $K_0$  and  $C$  used in the computation are given in table 1.

With these known values of controlling parameters, i.e., thermal conductivity ( $K(X, Z)$ ), heat generation ( $A(X, Z)$ ), mantle heat flow ( $Q_d$ ), length ( $L$ ) and depth ( $d$ ) of the model area, the subsurface temperature distribution is estimated by using the NISA program. More details of mathematical modeling procedure is given in Rai *et al* (2003) and Rai and Thiagarajan (2006).

## 5. Numerical results and discussion

The calculated surface heat flow values and isotherms along the profile are plotted in figure 3. The numerical results indicate that the Moho temperature varies between 470 and 500°C. The minimum value is in the northern part of the profile while the maximum value is in the southern part of the profile. The Moho temperature along this profile is found less than the Moho temperature along profiles I, II and III which is in the range of 500 to 580°C. The heat flow value varies between 45 to 47 mW/m<sup>2</sup> with the minimum value between Sendhwa and Tapti River and maximum between Jobat and Barwani. The surface heat flow values along this profile are found to be comparatively less than those surface heat flow values (46–49 mW/m<sup>2</sup>) along the other three profiles located in the east of BSF. The low surface heat flow between Narmada and Tapti is attributed to the presence of thick

volcanic layers which have a low heat generation value compared to the basement rocks. Though the heat flow values near to this profile are not available for comparison, however, these surface heat flow values are found in close agreement with heat flow values measured at Lonar (47 mW/m<sup>2</sup>), Satephal (47 mW/m<sup>2</sup>) and Singdad (49 mW/m<sup>2</sup>) in Yawatmal area near to the southern tips of profile II and Mandwa (48 mW/m<sup>2</sup>) located very near to the Multai–Pulgaon section of the profile III in Wardha area, and Mohapani (49 mW/m<sup>2</sup>) located in the northern tips of Satpura basin south of NSL (Gupta 1993; Roy and Rao 2000). Even south of Sindad in Ahmednagar region (around 75°E, 19–20°N) low heat flow values in the range of 35–45 mW/m<sup>2</sup> (mean 40 mW/m<sup>2</sup>) have been reported by Roy *et al* 1996.

The high surface heat flow values are also observed at Wadhona (71 mW/m<sup>2</sup>) and Palora (67 mW/m<sup>2</sup>) located very near to the Multai–Pulgaon section of profile II in Wardha region, and at Dattarampuram (59 mW/m<sup>2</sup>), Kurha Talni (66 mW/m<sup>2</sup>) and Loni (59 mW/m<sup>2</sup>) in the region of NW extension of Vaidarbha Nadi coalfield of Yawatmal district as shown in figure 1 (Roy and Rao 2000). Based on geological and geophysical investigations five major troughs are inferred in Wardha and Yawatmal regions starting from the Nagpur gravity low region towards south. These troughs are: (i) Kamthi, (ii) Umrer, (iii) Bhandar, (iv) Wardha–Pranhita, and (v) Vaidarbha Nadi coal field all extending in a northwesterly direction below the Deccan Traps (Joga Rao *et al* 1984; Chary 1993). Towards the southeastern side these troughs are connected to the probable northward extension of Pranhita–Godavari Gondwana graben through Wardha and Nagpur which abuts the NSL underneath Satpura basin (Qureshy *et al* 1968; Biswas 2003). These sites of high heat flow values are associated with gravity and resistivity lows indicating the presence of Gondwana sediments with fluids in general and coal seams in particular (Joga Rao *et al* 1984; Chary 1993; Gupta 1993). The presence of Gondwana sediments with coal seams has been confirmed by boreholes data drilled near Hingna (79°E, 21°5′N) in Umrer trough, and Astona (78°45′E, 20°15′N) in northwesterly extension of Rajur coal field of Wardha–Pranhita trough. On the other hand, sites of heat flow values of the order of 49 mW/m<sup>2</sup> and less at Lonar, Satephal, Singdad, and Mandwa are located at boundaries of two adjacent troughs. These boundaries are characterized with less resistivity, indicating that these values are free from the effect of hydrothermal circulation and represent the conductive surface heat flow value for NSL region in central India.

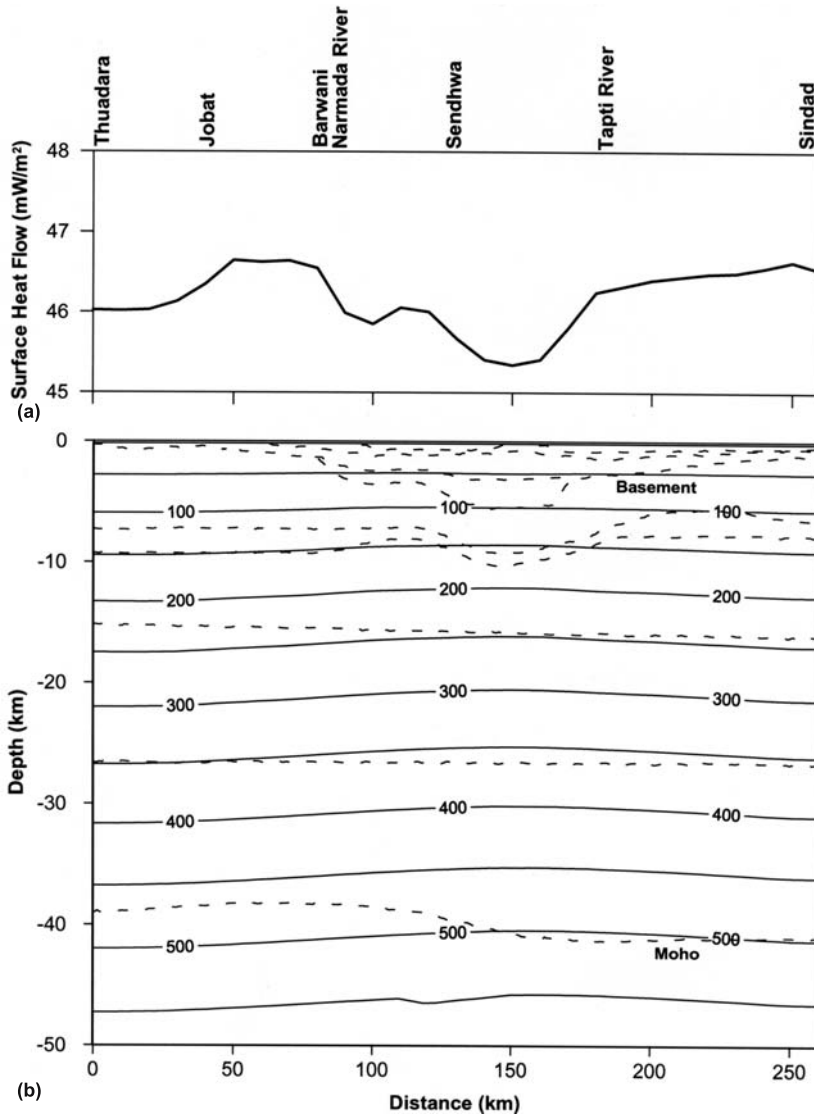


Figure 3. Calculated (a) surface heat flow values and (b) isotherms ( $^{\circ}\text{C}$ ) (solid line) along with crustal structure (dotted line) along the profile.

The high surface heat flow values are also observed at Damua ( $62\text{ mW/m}^2$ ), Parasia ( $90\text{ mW/m}^2$ ) and Mandla ( $96\text{ mW/m}^2$ ) in Pench Kanhan Valley coal field located south of NSL in Satpura basin. The presence of hot fluids in the sedimentary rocks may give rise to the high heat flow value. Basic dykes and sills which have an affinity to Deccan Traps and several faults such as Parasia fault do exist in the Satpura basin (Datta 1993; Gupta 1993; Biswas 2003). These dykes and faults control the water movement in the basin. The dyke, if located towards the down-dip side of the basin, would act as a barrier and assist upward flow of deep circulating hot water. Faults also work as conduit for the upward movement of hot water. This process would result in raising the magnitude of surface heat flow values at the above-mentioned sites similar to that reported from Ashwaraopet

( $104\text{ mW/m}^2$ ) and Chintalapudi ( $92\text{ mW/m}^2$ ) in Godavari Valley (Rao *et al* 1970; Gupta 1993). This fact is supported by the observation of heat flow rise due to hydrothermal circulation across a layer with  $10^{\circ}\text{C}$  temperature difference which would contribute about  $40\text{ mW/m}^2$  to the surface heat flow density value (Lachenbruch and Sass 1977). Based on the results of rheological modeling, Manglik and Singh (2002) have also suggested that occurrence of deep crustal events at  $\sim 35\text{--}37\text{ km}$  depth such as Jabalpur and Satpura earthquakes require the surface heat flow value lower than  $48\text{ mW/m}^2$ . There are several other examples of deep earthquakes ( $25\text{--}40\text{ km}$ ) such as Saguenay (1988) in east Canada; Soleberg (1938) in Sweden; Manaus (1963) in Brazil and Brome (1979) in Australia which occurred in the areas characterized with  $< 50\text{ mW/m}^2$  heat flow value (Chen 1988).

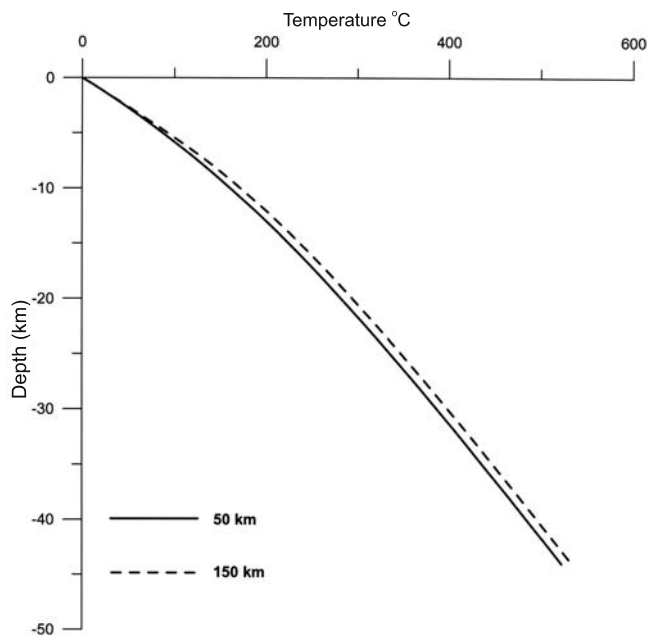


Figure 4. Calculated temperature–depth distribution at 50 km distance (solid line) and at 150 km distance (dotted line).

The Curie depth is found to vary between 46 and 47 km and is in close agreement with the earlier reported Curie depth (46 km) estimated from the analysis of MAGSAT data along the profile P208 which passes through the NSL region under consideration (Agarwal *et al* 1992). The Curie depth is  $\sim 10$  km below the Moho in the north of Sendhwa and is  $\sim 5$  km below the Moho in the south of Sendhwa.

Temperature of B/D transition when compared with the maximum depth of seismicity gives a probable estimate of prevailing temperature at that depth (Manglik and Singh 2002). The most probable value adopted for the focal depth of Satpura range earthquake of 1938 is 40 km (Mukherjee 1942). Experimental results have shown that quartz, feldspar and olivine become ductile at 300–450°C and 700–750°C respectively, for natural strain rates (Meissner and Strehleau 1982; Chen and Molnar 1983). Therefore, temperature of B/D transition for mafic rock should be between 450 and 700°C. It is evident from figure 3 that the south of Sendhwa, in the region of Satpura range, the focal depth (40 km) of Satpura earthquake coincides with the 500°C isotherm which may be regarded as temperature of B/D transition in the mafic lower crust. This depth incidentally coincides with the Moho depth. However, this is not the case north of BSF where Moho is found to be 2–3 km above the 500°C isotherm.

The calculated temperature–depth profiles at 50 km distance outside the graben near Jobat and

at 150 km distance within the graben between Sendhwa and Tapti River are plotted in figure 4. The figure illustrates the nature of variation in temperature distribution at both locations. The temperature gradient up to a characteristic depth (11.5 km) is around 16°C/km and below up to the Moho is around 10.5°C/km.

There is always a possibility of uncertainty in the estimation of heat generation and thermal conductivity which is likely to affect the temperature field. By considering a 1-D homogeneous crustal model of 35 km thickness having surface heat generation =  $2.0 \mu\text{Wm}^{-3}$ , mantle heat flow =  $25 \text{ mW/m}^2$ , mean value of  $K = 2.5 \text{ W/mK}$  with  $0.25 \text{ W/mK}$  standard deviation (S.D.). Srivastava and Singh (1998) have computed mean temperature equal to 419°C with 44°C S.D. at 35 km depth which amounts to  $\sim 10\%$  deviation in the temperature field at Moho. Similarly by considering another 1-D crustal model of 35 km thickness having mantle heat flow =  $23 \text{ mW/m}^2$ ,  $K = 3.0 \text{ W/mK}$ , mean surface heat generation =  $1.0 \mu\text{Wm}^{-3}$  with  $0.5 \mu\text{Wm}^{-3}$  S.D., they have calculated Moho mean temperature around 482°C with 102°C S.D. which amounts to  $\sim 21\%$  deviation in the temperature field (Srivastava and Singh 1999). These results are discussed here to highlight the effects of uncertainty in heat generation and conductivity values on the temperature field.

## 6. Conclusions

The numerical results indicate that the surface heat flow value in the western part of NSL along Thuadara–Sindad profile varies between 45 and  $47 \text{ mW/m}^2$  and is in close agreement with heat flow values measured at some places such as Lonar, Satephal, Singdad, Mandwa and Mohapani located south of NSL. The Curie depth is  $\sim 46$ – $47$  km which is in close agreement with the Curie depth (46 km) estimated from the analysis of MAGSAT data. The calculated surface heat flow value is found to be much less than the earlier suggested values by Ravi Shanker (1988) in the range of 70– $100 \text{ mW/m}^2$  for NSL region. His values are based on silica content in groundwater and mostly in the zones of upwelling thermal water. In such areas convective transfer of heat is predominant and any heat flow measurement is not reflective of conductive regional heat flow. Numerical results indicate that for surface heat flow value of the order of  $70 \text{ mW/m}^2$ , the mantle heat flow has to be increased to  $47 \text{ mW/m}^2$ . But this value of mantle heat flow gives Curie depths at  $\sim 17$  km. This value does not match with the Curie depth

(~46–47 km) estimated from the MAGSAT data. The Moho temperature varies between 470 and 500°C, which is less than the Moho temperature reported for Cambay basin (~900°C) and Godavari graben/basin (~620°C). This indicates that the lower crust in this region is cool which in turn makes it of brittle nature amenable to the occurrence of deep focused earthquakes such as Satpura (1938).

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